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Figures

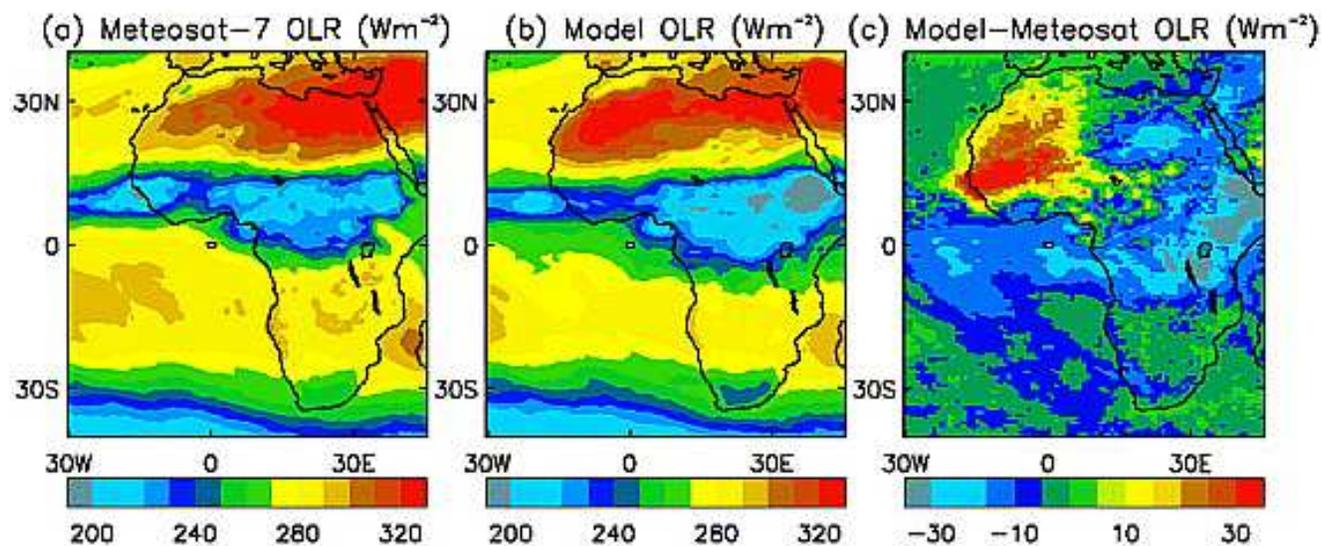


Figure 1. The July 2003 monthly mean for (a) OLR_{Met7} , (b) $\text{OLR}_{\text{model}}$, and (c) $\text{OLR}_{\text{model}} - \text{OLR}_{\text{Met7}}$. The monthly mean consists of the average of the monthly mean of the OLR diagnosed at 0000 UTC, 6000 UTC, 1200 UTC, and 1800 UTC. Units are Wm^{-2} .

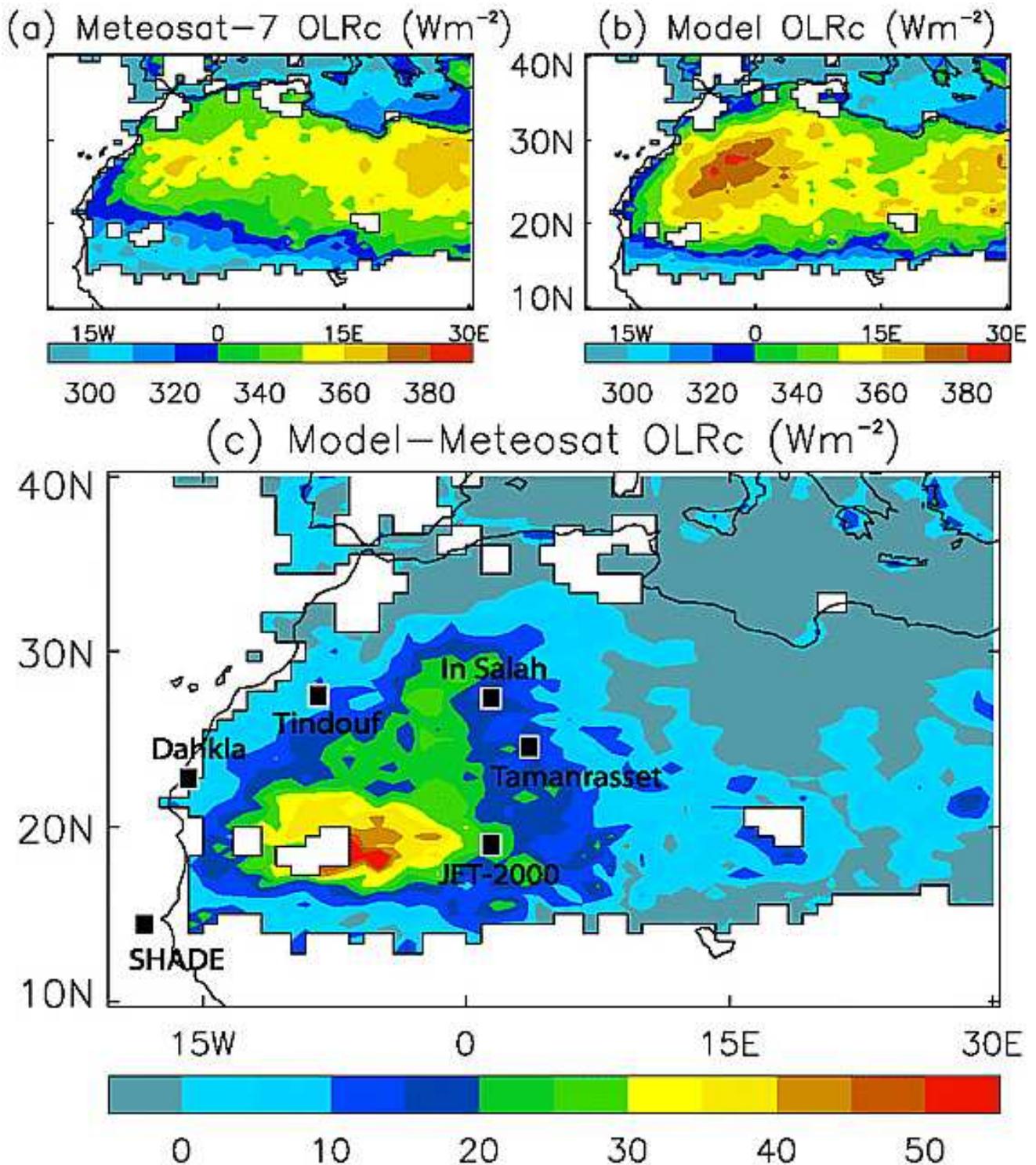


Figure 2. The July 2003 monthly mean of the 1200 UTC clear sky for (a) $\text{OLRc}_{\text{Met7}}$, (b) $\text{OLRc}_{\text{model}}$, and (c) $\text{OLRc}_{\text{model}} - \text{OLRc}_{\text{Met7}}$. Units are Wm^{-2} . Areas shown in white indicate missing/cloudy data. The approximate positions of the radiosonde stations, measurement campaigns, and AERONET sites referred to in the text are also shown.

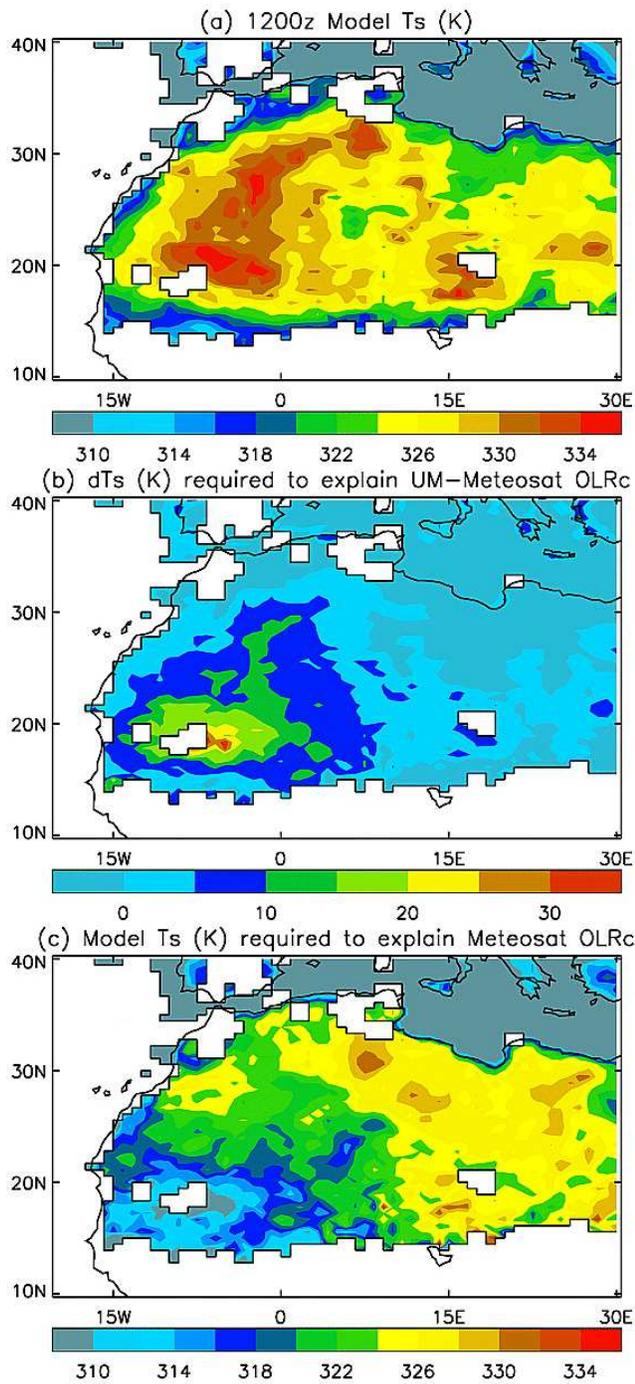


Figure 4. (a) Surface skin temperature, T_s , from the NWP model. (b) The reduction in the model temperature required to explain the difference between $OLRc_{model}$ and $OLRc_{Met7}$. (c) The model T_s required so that $OLRc_{model}$ fits $OLRc_{Met7}$. Units K.

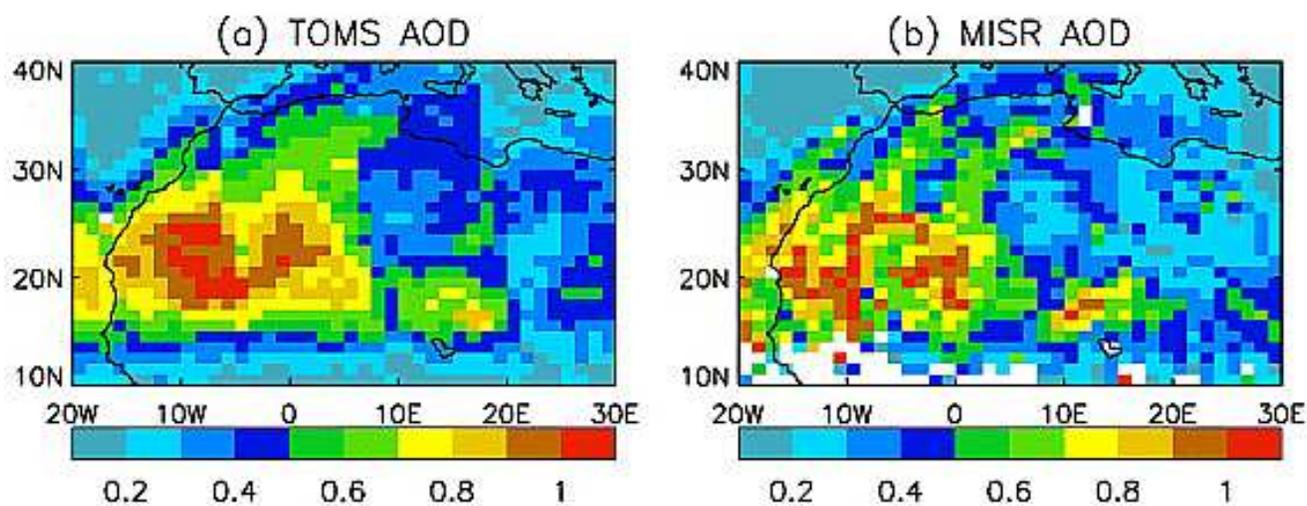


Figure 8. The $\tau_{\text{aer}0.55}$ derived for July 2003 for (a) TOMS and (b) MISR. In deriving the $\tau_{\text{aer}0.55}$ from TOMS, the TOMS AI is related to $\tau_{\text{aer}0.44}$ via [equation \(3\)](#), and the wavelength dependence of k_e shown in [Figure 4a](#) is used to derive $\tau_{\text{aer}0.55}$ via [equation \(4\)](#). The $\tau_{\text{aer}0.55}$ may be related to $\tau_{\text{aer}10}$ by [equations \(5\)](#) and [\(6\)](#).

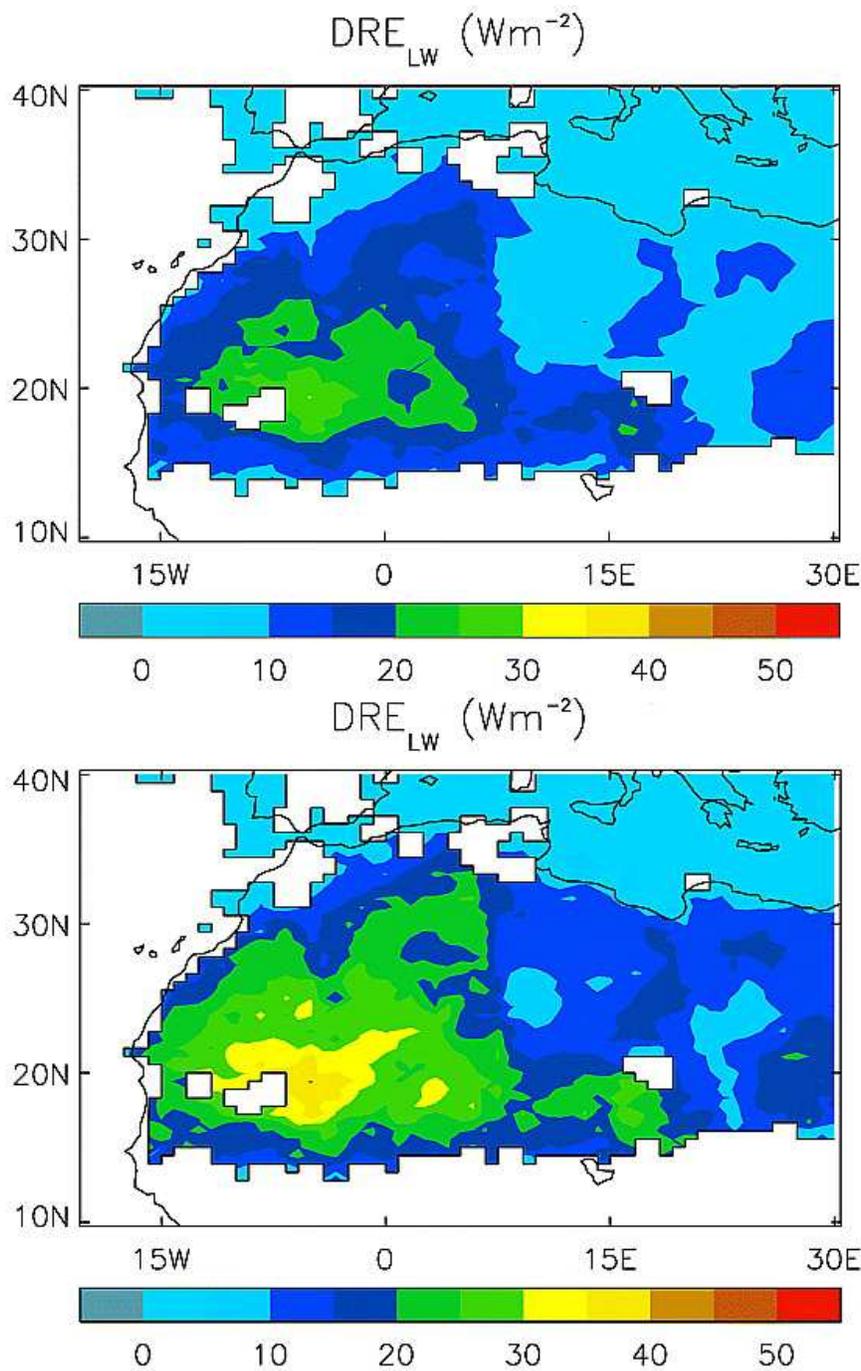


Figure 11. DRE_{LW} due to the radiative effect of mineral dust (Wm^{-2}) for refractive indices from (a) [Fouquart et al. \[1987\]](#) and (b) [Volz \[1973\]](#).

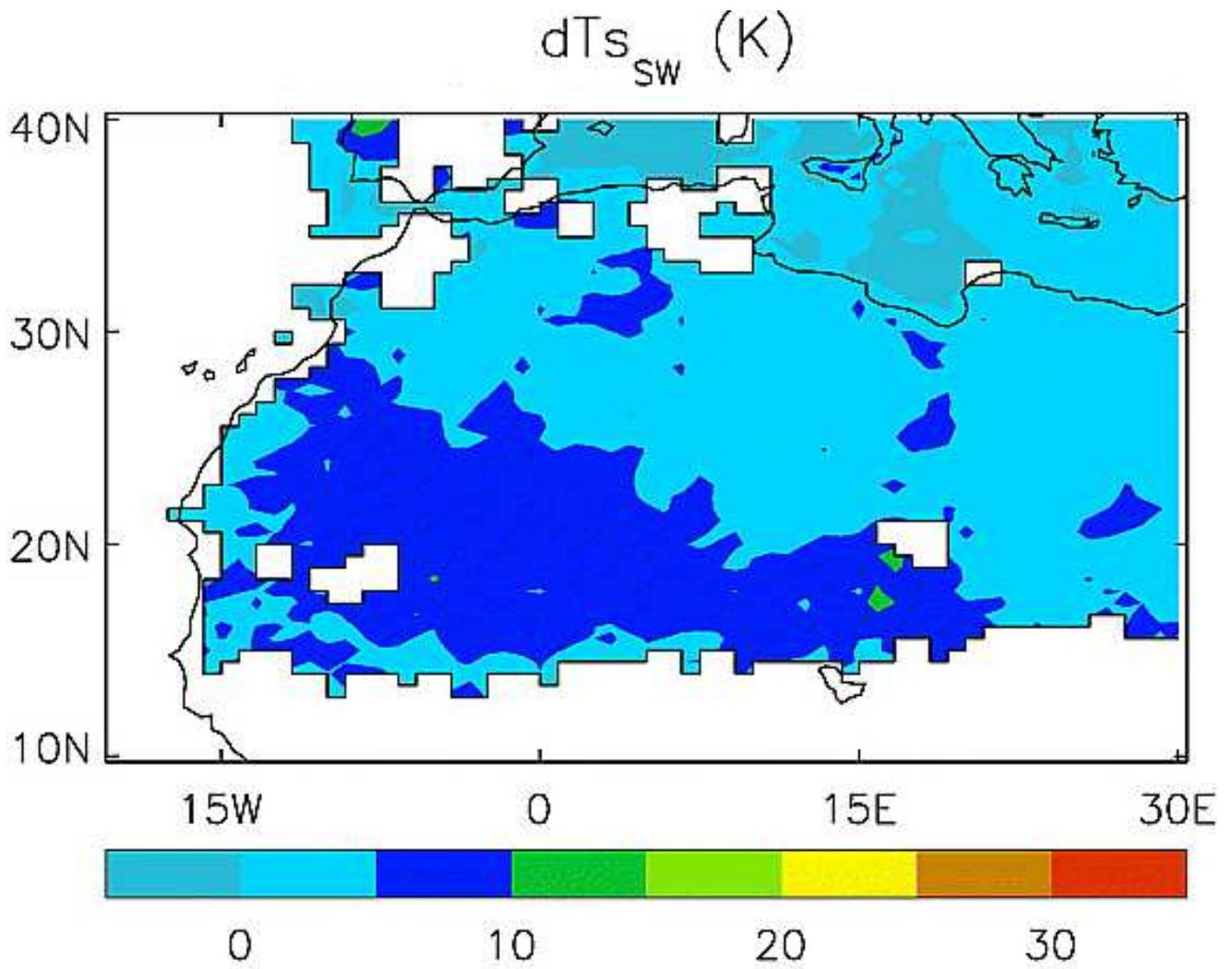


Figure 13. The change in model T_s (K) due to the reduction in SW_{surf} .

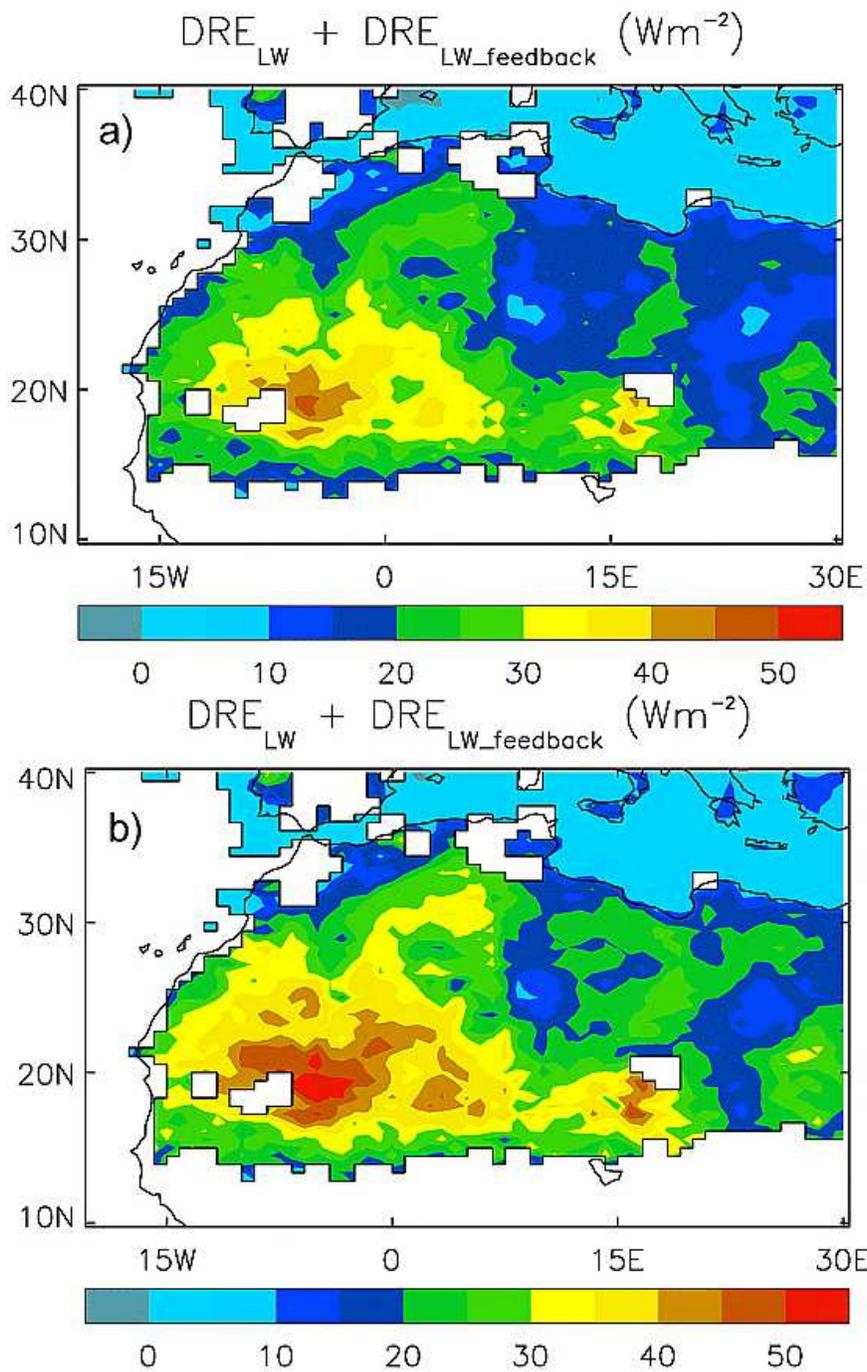


Figure 14. $DRE_{LW} + DRE_{LW_feedback} \text{ (Wm}^{-2}\text{)}$, which represents our best estimate of the radiative effect of mineral dust in the terrestrial spectrum, (a) using the refractive indices of [Fouquart et al. \[1987\]](#) and (b) using the refractive indices of [Volz \[1973\]](#). The refractive indices of [WCP \[1986\]](#) are used in the solar spectral range as described in the text.

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